Ergodicity in data assimilation methods

David Kelly

Andy Majda Xin Tong

Courant Institute New York University New York NY www.dtbkelly.com

April 15, 2016

ETH Zurich

David Kelly (CIMS)

Data assimilation

April 15, 2016 1 / 29

Part I: What is data assimilation?

David Kelly (CIMS)

Data assimilation

April 15, 2016 2 / 29

What is data assimilation?

Suppose u satisfies an evolution equation

$$\frac{d \mathbf{u}}{dt} = F(\mathbf{u})$$

with some **unknown** initial condition $u_0 \sim \mu_0$.

There is a true trajectory of u that is producing *partial*, *noisy* observations at times t = h, 2h, ..., nh:

$$\mathbf{y}_n = H\mathbf{u}_n + \boldsymbol{\xi}_n$$

where *H* is a linear operator (think low rank projection), $u_n = u(nh)$, and $\xi_n \sim N(0, \Gamma)$ iid.

The aim of **data assimilation** is to characterize the conditional distribution of u_n given the observations $\{y_1, \ldots, y_n\}$

David Kelly (CIMS)

The conditional distribution is updated via the **filtering cycle**.

David Kelly (CIMS)

Data assimilation

April 15, 2016 4 / 29

Illustration (Initialization)

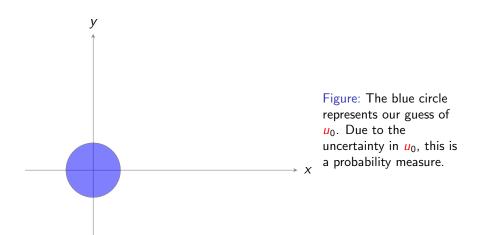


Illustration (Forecast step)

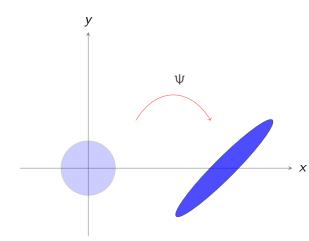


Figure: Apply the time h flow map Ψ . This produces a new probability measure which is our forecasted estimate of u_1 . This is called the forecast step.

Illustration (Make an observation)

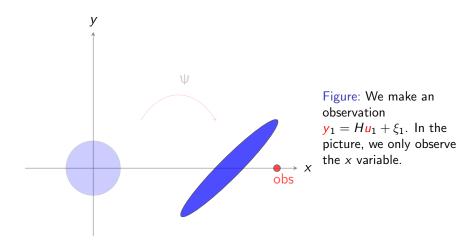


Illustration (Analysis step)

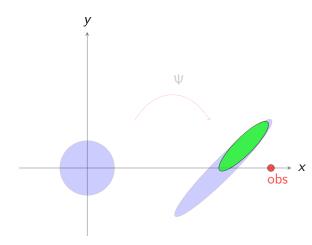


Figure: Using Bayes formula we compute the conditional distribution of $u_1 | y_1$. This new measure (called the posterior) is the new estimate of u_1 . The uncertainty of the estimate is reduced by incorporating the observation. The forecast distribution steers the update from the observation.

Bayes' formula filtering update

Let $\mathbf{Y}_n = \{\mathbf{y}_0, \mathbf{y}_1, \dots, \mathbf{y}_n\}$. We want to compute the conditional density $\mathbf{P}(u_{n+1}|\mathbf{Y}_{n+1})$, using $\mathbf{P}(u_n|\mathbf{Y}_n)$ and \mathbf{y}_{n+1} .

By Bayes' formula, we have

$$\mathsf{P}(\underline{u}_{n+1}|\underline{Y}_{n+1}) = \mathsf{P}(\underline{u}_{n+1}|\underline{Y}_n, \underline{y}_{n+1}) \propto \mathsf{P}(\underline{y}_{n+1}|\underline{u}_{n+1})\mathsf{P}(\underline{u}_{n+1}|\underline{Y}_n)$$

But we need to compute the integral

$$\mathbf{P}(\boldsymbol{u}_{n+1}|\boldsymbol{Y}_n) = \int \mathbf{P}(\boldsymbol{u}_{n+1}|\boldsymbol{Y}_n, \boldsymbol{u}_n) \mathbf{P}(\boldsymbol{u}_n|\boldsymbol{Y}_n) d\boldsymbol{u}_n .$$

The 'optimal filtering' framework is **impractical** for high dimensional models, as the integrals are impossible to compute and densities impossible to store. In **numerical weather prediction**, we have $O(10^9)$ variables for ocean-atmosphere models (discretized PDEs). The **Ensemble Kalman filter** (EnKF) is a low dimensional representation, empirical approximation of the posterior measure $P(u_n | Y_n)$.

EnKF builds on the idea of a Kalman filter.

The Kalman Filter

For a linear model $u_{n+1} = Mu_n + \eta_{n+1}$, the Bayesian integral is Gaussian and can be computed explicitly. The conditional density is proportional to

$$\exp\left(-\frac{1}{2}|\Gamma^{-1/2}(\mathbf{y}_{n+1}-Hu)|^2-\frac{1}{2}|\widehat{C}_1^{-1/2}(u-\widehat{\mathbf{m}}_{n+1})|^2\right)$$

where $(\widehat{m}_{n+1}, \widehat{c}_{n+1})$ is the **forecast** mean and covariance. We obtain

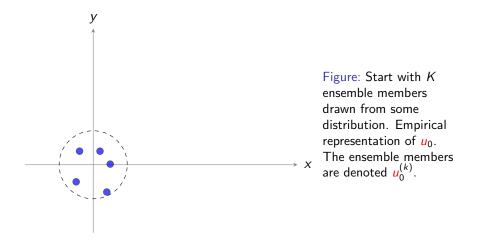
$$m_{n+1} = (1 - K_{n+1}H)\widehat{m}_n + K_{n+1}y_{n+1}$$

$$C_{n+1} = (I - K_{n+1}H)\widehat{C}_{n+1},$$

where $K_{n+1} = \widehat{C}_{n+1}H^T(\Gamma + H\widehat{C}_{n+1}H^T)^{-1}$ is the **Kalman gain**. The procedure of updating $(m_n, C_n) \mapsto (m_{n+1}, C_{n+1})$ is known as the **Kalman filter**.

David Kelly (CIMS)

Ensemble Kalman filter (Evensen 94)



Only KN numbers are stored. Better than Kalman if K < N.

David Kelly (CIMS)

Ensemble Kalman filter (Forecast step)

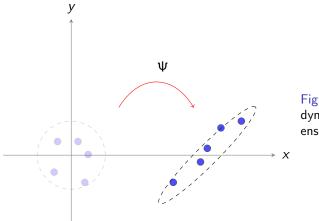
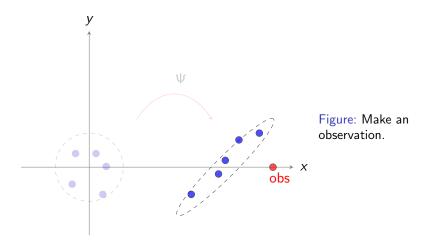


Figure: Apply the dynamics Ψ to each ensemble member.

Ensemble Kalman filter (Make obs)



Ensemble Kalman filter (Analysis step)

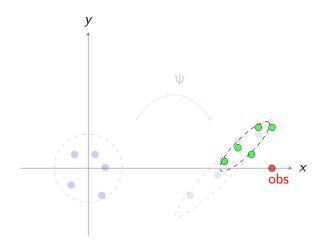


Figure: Approximate the forecast distribution with a Gaussian. Fit the Gaussian using the empirical statistics of the ensemble.

How to implement the Gaussian approximation

The posterior is *approximately sampled* using the **Randomized Maximum Likelihood** (RML) method: Draw the sample $u_{n+1}^{(k)}$ by minimizing the functional

$$\frac{1}{2}|\Gamma^{-1/2}(\mathbf{y}_{n+1}^{(k)}-Hu)|^2+\frac{1}{2}|\widehat{\boldsymbol{C}}_1^{-1/2}(u-\Psi^{(k)}(\boldsymbol{u}_n^{(k)}))|^2$$

where $y_{n+1}^{(k)} = y_{n+1} + \xi_{n+1}^{(k)}$ is a perturbed observation.

In the linear case $\Psi(u_n) = Mu_n + \eta_n$, this produces iid Gaussian samples with mean and covariance satisfying the Kalman update equations, with \hat{C} in place of the true forecast covariance.

We end up with
$$m{u}_{n+1}^{(k)} = (1 - m{K}_{n+1} H) \Psi^{(k)}(m{u}_n^{(k)}) + m{K}_{n+1} H m{y}_{n+1}^{(k)}.$$

David Kelly (CIMS)

Ensemble Kalman filter (Perturb obs)

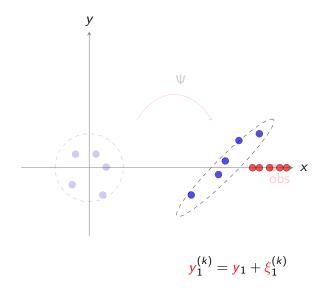


Figure: Turn the observation into *K* artificial observations by perturbing by the same source of observational noise.

Ensemble Kalman filter (Analysis step)

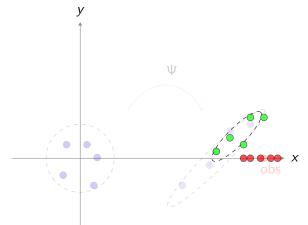


Figure: Update each member using the Kalman update formula. The Kalman gain K_1 is computed using the ensemble covariance.

$$\begin{aligned} \mathbf{u}_{1}^{(k)} &= (1 - K_{1}H)\Psi(\mathbf{u}_{0}^{(k)}) + K_{1}H\mathbf{y}_{1}^{(k)} \quad K_{1} = \widehat{C}_{1}H^{T}(\Gamma + H\widehat{C}_{1}H^{T})^{-1} \\ & \widehat{C}_{1} = \frac{1}{K - 1}\sum_{k=1}^{K}(\Psi(\mathbf{u}_{0}^{(k)}) - \widehat{m}_{n+1})(\Psi(\mathbf{u}_{0}^{(k)}) - \widehat{m}_{n+1})^{T} \\ & \text{David Kelly (CIMS)} \quad April 15, 2016 \quad 12 / 29 \end{aligned}$$

What are we interested in understanding?

1– Stability with respect to initializations (ergodicity): Statistics of $(u_n, u_n^{(1)}, \dots, u_n^{(K)})$ converge to invariant statistics as $n \to \infty$.

- 2 Accuracy of the approximation: Law of $(u_n^{(1)}, \ldots, u_n^{(K)})|Y_n$ approximates posterior as $n \to \infty$.
 - **Rmk**. All results are in the *K* fixed regime.

Today we focus on geometric ergodicity for the **signal-filter** process $(u_n, u_n^{(1)}, \dots, u_n^{(K)}).$

The theoretical framework

A Markov chain $\{X_n\}_{n\in\mathbb{N}}$ on a state space \mathcal{X} is called **geometrically** ergodic if it has a unique invariant measure π and for any initialization X_0 we have

$$\sup_{|f|\leq 1} \left| \mathbf{E}_{\mathbf{x}_0} f(\mathbf{X}_n) - \int f(\mathbf{x}) \pi(d\mathbf{x}) \right| \leq C(\mathbf{x}_0) r^n$$

for some $r \in (0,1)$ and any m-ble bdd f.

We use the **coupling** approach: Let X'_n and X''_n be two copies of X_n , such that $X'_0 = x_0$ and $X''_0 \sim \pi$, and are coupled in such a way that $X'_0 = X''_n$ for $n \geq T$, where T is the first hitting time $X'_T = X''_T$. Then we have

$$\|P^n\delta_{\mathbf{x}_0}-\pi\|_{TV}\leq 2\mathbf{P}(\mathbf{T}>n)$$

The **Doeblin** / **Meyn-Tweedie** approach is to verify two assumptions that guarantee the coupling can be constructed with $P(T > n) \leq r^n$:

- 1- Lyapunov function / Energy dissipation: $\mathbf{E}_n |X_{n+1}|^2 \le \alpha |X_n|^2 + \beta$ with $\alpha \in (0, 1)$, $\beta > 0$.
- **2-** Minorization: Find compact $C \subset \mathcal{X}$, measure ν supported on C, $\kappa > 0$ such that $P(x, A) \ge \kappa \nu(A)$ for all $x \in C$, $A \subset \mathcal{X}$.

Minorization is inherited from model

Smooth densities on $C \Rightarrow$ minorization. For the signal-filter process $(u_n, u_n^{(1)}, \dots, u_n^{(K)})$ we can write the Markov chain as a composition of two maps:

$$\begin{array}{cccc} (u_{n}, u_{n}^{(1)}, \dots, u_{n}^{(K)}) & \to & (u_{n+1}, u_{n+1}^{(1)}, \dots, u_{n+1}^{(K)}) \\ & & & \uparrow \\ & & & (u_{n+1}, \Psi^{(1)}(u_{n}^{(1)}), \dots, \Psi^{(K)}(u_{n}^{(K)}), y_{n}^{(1)}, \dots, y_{n}^{(K)}) \end{array}$$

The first step is a nice Markov kernel with smooth densities (by assumptions on model and observational noise). The second is a deterministic map.

We only need to check that deterministic map preserves the densities (just near one chosen point y^*).

David Kelly (CIMS)

Lyapunov function: Inheriting an energy principle

Suppose we know the model satisfies an energy principle

$$\mathbf{E}_n |\Psi(\boldsymbol{u})|^2 \le \alpha |\boldsymbol{u}|^2 + \beta$$

for $\alpha \in (0, 1), \beta > 0$. Does the filter inherit the energy principle?

$$\mathbf{E}_n |\mathbf{u}_{n+1}^{(k)}|^2 \le \alpha' |\mathbf{u}_n^{(k)}|^2 + \beta'$$

Only observable energy is inherited (Tong, Majda, K. 16)

We can show that if the model satisfies an observable energy criterion:

$$\mathbf{E}_n |H\mathbf{u}_{n+1}|^2 \le \alpha |H\mathbf{u}_n|^2 + \beta$$

then the filter will too:

$$\mathbf{E}_{n}|H\mathbf{u}_{n+1}^{(k)}|^{2} \leq \alpha'|H\mathbf{u}_{n}^{(k)}|^{2} + \beta'$$

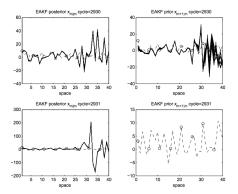
The observable energy criterion holds when there is 'no interaction' between observed and unobserved variables at high energies.

But when H is not full rank this does not give a Lyapunov function, so **no ergodicity**, only boundedness of observed components.

David Kelly (CIMS)

Catastrophic filter divergence

Lorenz-96: $\mathbf{u}_j = (\mathbf{u}_{j+1} - \mathbf{u}_{j-2})\mathbf{u}_{j-1} - \mathbf{u}_j + F$ with $j = 1, \dots, 40$. Periodic BCs. Observe every fifth node. (Harlim-Majda 10, Gottwald-Majda 12)



True solution in a bounded set, but filter **blows up** to machine infinity in finite time!

David Kelly (CIMS)

For complicated models, only heuristic arguments offered as explanation.

Can we **prove** it for a simpler constructive model?

David Kelly (CIMS)

The rotate-and-lock map (K., Majda, Tong. PNAS 15.)

The model $\Psi : \mathbb{R}^2 \to \mathbb{R}^2$ is a composition of two maps $\Psi(x, y) = \Psi_{lock}(\Psi_{rot}(x, y))$ where

$$\Psi_{rot}(x,y) = \begin{pmatrix} \rho \cos \theta & -\rho \sin \theta \\ \rho \sin \theta & \rho \cos \theta \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix}$$

and $\Psi_{\textit{lock}}$ rounds the input to the nearest point in the grid

$$\mathcal{G} = \{(m,(2n+1)arepsilon)\in\mathbb{R}^2:m,n\in\mathbb{Z}\}$$
 .

It is easy to show that this model has an energy dissipation principle:

$$|\Psi(x,y)|^2 \le \alpha |(x,y)|^2 + \beta$$

for $\alpha \in (0, 1)$ and $\beta > 0$.

David Kelly (CIMS)

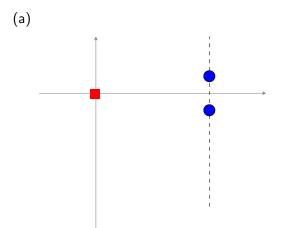
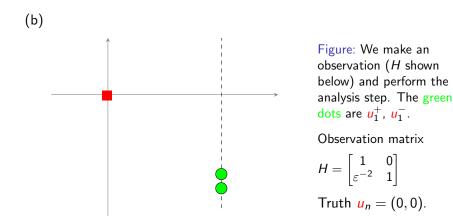


Figure: The red square is the trajectory $u_n = 0$. The blue dots are the positions of the forecast ensemble $\Psi(u_0^+)$, $\Psi(u_0^-)$. Given the locking mechanism in Ψ , this is a natural configuration.



The filter is certain that the x-coordinate is \hat{x} (the dashed line). The filter thinks the observation must be $(\hat{x}, \varepsilon^{-2}\hat{x} + u_{1,y})$, but it is actually (0,0) + noise. The filter concludes that $u_{1,y} \approx -\varepsilon^{-2}\hat{x}$.

David Kelly (CIMS)

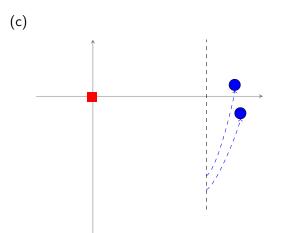


Figure: Beginning the next assimilation step. Apply Ψ_{rot} to the ensemble (blue dots)

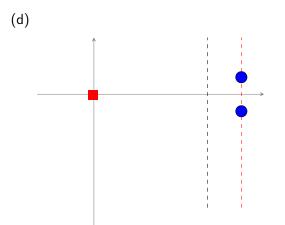


Figure: Apply Ψ_{lock} . The blue dots are the forecast ensemble $\Psi(\boldsymbol{u}_1^+), \Psi(\boldsymbol{u}_1^-)$. Exact same as frame 1, but higher energy orbit. The cycle repeats leading to **exponential growth**. Theorem (K., Majda, Tong 15 PNAS) For any N > 0 and any $p \in (0, 1)$ there exists a choice of parameters such that

$$\mathsf{P}\left(|oldsymbol{u}_n^{(k)}|\geq M_n ext{ for all } n\leq N
ight)\geq 1-p$$

where M_n is an exponentially growing sequence.

ie - The filter can be made to grow exponentially for an arbitrarily long time with an arbitrarily high probability.

Ensemble alignment can cause EnKF to gain energy, even when the model dissipates energy.

Can we get around the problem by **tweaking** the algorithm?

Adaptive Covariance Inflation (Tong, Majda, K. 15) We modify algorithm by introducing a **covariance inflation** :

$$\widehat{\boldsymbol{C}}_{n+1}\mapsto \widehat{\boldsymbol{C}}_{n+1}+\lambda_{n+1}\boldsymbol{I}$$

where

$$\lambda_{n+1} \propto \Theta_{n+1} \mathbf{1}(\Theta_{n+1} > \Lambda)$$
$$\Theta_{n+1} = \sqrt{\frac{1}{K} \sum_{k=1}^{K} |\mathbf{y}_{n+1}^{(k)} - H\Psi(\mathbf{u}_{n}^{(k)})|^2}$$

and Λ is some constant threshold. If the predictions are near the observations, then there is no inflation.

Thm. The modified EnKF inherits an energy principle from the model.

$$\mathbf{E}_{x}|\Psi(x)|^{2} \leq \alpha|x|^{2} + \beta \Rightarrow \mathbf{E}_{n}|\boldsymbol{u}_{n+1}^{(k)}|^{2} \leq \alpha'|\boldsymbol{u}_{n}^{(k)}|^{2} + \beta'$$

Consequently, the modified EnKF is signal-filter ergodic.

David Kelly (CIMS)

Adaptive inflation schemes allows us to use **cheap** integration schemes in the forecast dynamics. These would usually lead to numerical blow-up.

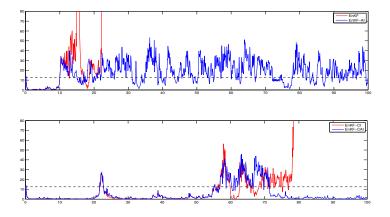


Figure: RMS error for EnKF on 5d Lorenz-96 with sparse obs (1 node), strong turbulence regime. Euler method with course step size. Lower panel has additional constant inflation which helps accuracy.

Applicable to more sophisticated geophysical models, such as 2-layer QG with course graining (Lee, Majda 16').

David Kelly (CIMS)

References

1 - D. Kelly, K. Law & A. Stuart. *Well-Posedness And Accuracy Of The Ensemble Kalman Filter In Discrete And Continuous Time.* **Nonlinearity** (2014).

2 - D. Kelly, A. Majda & X. Tong. *Concrete ensemble Kalman filters with rigorous catastrophic filter divergence*. **Proc. Nat. Acad. Sci.** (2015).

3 - X. Tong, A. Majda & D. Kelly. *Nonlinear stability and ergodicity of ensemble based Kalman filters*. **Nonlinearity** (2016).

4 - X. Tong, A. Majda & D. Kelly. *Nonlinear stability of the ensemble Kalman filter with adaptive covariance inflation*. To appear in **Comm. Math. Sci.** (2016).

All my slides are on my website (www.dtbkelly.com) Thank you!

David Kelly (CIMS)