### EnKF and Catastrophic filter divergence

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## Talk outline

- **1**. What is EnKF?
- 2. What is known about EnKF?
- **3**. How can we use stochastic analysis to better understand EnKF?

## The filtering problem

We have a deterministic model

$$rac{d\mathbf{v}}{dt} = F(\mathbf{v}) \quad ext{with } \mathbf{v}_0 \sim N(m_0, C_0) \; .$$

We will denote  $v(t) = \Psi_t(v_0)$ . Think of this as very high dimensional and nonlinear.

We want to estimate  $v_j = v(jh)$  for some h > 0 and j = 0, 1, ..., J given the observations

$$y_j = Hv_j + \xi_j$$
 for  $\xi_j$  iid  $N(0, \Gamma)$ .

## We can write down the conditional density using **Bayes' formula** ...

But for high dimensional nonlinear systems it's horrible.

## Alternatively, we can use **EnKF** to draw **approximate samples** from the posterior.

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## For linear models, one can draw samples, using the Randomized Maximum Likelihood method.

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### RML method

Let  $u \sim N(\hat{m}, \hat{C})$  and  $\eta \sim N(0, \Gamma)$ . We make an observation  $\mathbf{v} = H\mathbf{u} + \eta$ .

We want the conditional distribution of u|y. This is called an **inverse** problem.

RML takes a sample

$$\{\widehat{u}^{(1)},\ldots,\widehat{u}^{(K)}\}\sim N(\widehat{m},\widehat{C})$$

and turns them into a sample

 $\{\mathbf{u}^{(1)},\ldots,\mathbf{u}^{(K)}\}\sim\mathbf{u}|\mathbf{y}|$ 

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#### RML method: How does it work?

Along with the prior sample  $\{\hat{u}^{(1)}, \ldots, \hat{u}^{(\kappa)}\}\)$ , we create **artificial observations**  $\{y^{(1)}, \ldots, y^{(\kappa)}\}\)$  where

$$\mathbf{y}^{(k)} = \mathbf{y} + \eta^{(k)}$$
 where  $\eta^{(k)} \sim \mathsf{N}(0, \mathsf{\Gamma})$  i.i.d

Then define  $u^{(k)}$  using the **Bayes formula** update, with  $(\hat{u}^{(k)}, y^{(k)})$ 

$$u^{(k)} = \widehat{u}^{(k)} + G(\widehat{u})(y^{(k)} - H\widehat{u}^{(k)}).$$

Where the "Kalman Gain"  $G(\hat{u})$  is computing using the covariance of the prior  $\hat{u}$ .

The set  $\{u^{(1)}, \ldots, u^{(K)}\}$  are exact samples from u|y.

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# EnKF uses the same method, but with an **approximation** of the covariance in the Kalman gain.

## The set-up for EnKF

Suppose we are given the ensemble  $\{u_j^{(1)}, \ldots, u_j^{(K)}\}$  at time *j*. For each ensemble member, we create an **artificial observation** 

$$y_{j+1}^{(k)} = y_{j+1} + \xi_{j+1}^{(k)}$$
,  $\xi_{j+1}^{(k)}$  iid  $N(0, \Gamma)$ .

We update each particle using the Kalman update

$$u_{j+1}^{(k)} = \Psi_h(u_j^{(k)}) + G(u_j) \left( y_{j+1}^{(k)} - H \Psi_h(u_j^{(k)}) \right) ,$$

where  $G(u_j)$  is the Kalman gain computed using the forecasted ensemble covariance

$$\widehat{C}_{j+1} = \frac{1}{K} \sum_{k=1}^{K} (\Psi_h(\boldsymbol{u}_j^{(k)}) - \overline{\Psi_h(\boldsymbol{u}_j)})^T (\Psi_h(\boldsymbol{u}_j^{(k)}) - \overline{\Psi_h(\boldsymbol{u}_j)}) .$$

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## There aren't many **theorems** about EnKF.

## Ideally, we would like a theorem about **long time behaviour** of the filter for a finite ensemble size.

## Filter divergence

In certain situations, it has been observed  $(\star)$  that the ensemble can **blow-up** (ie. reach machine-infinity) in **finite time**, even when the model has nice bounded solutions.

This is known as catastrophic filter divergence.

We would like to investigate whether this has a **dynamical justification** or if it is simply a **numerical artefact**.

★ Harlim, Majda (2010), Gottwald (2011), Gottwald, Majda (2013).

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## Assumptions on the dynamics

We make a dissipativity assumption on the model. Namely that

$$\frac{d\mathbf{v}}{dt} + A\mathbf{v} + B(\mathbf{v}, \mathbf{v}) = f$$

with A linear elliptic and B bilinear, satisfying certain estimates and symmetries.

This guarantees uniformly bounded solutions.

Eg. 2d-Navier-Stokes, Lorenz-63, Lorenz-96.

### Discrete time results

For a fixed observation frequency h > 0 we can prove

Theorem (AS,DK) If  $H = \Gamma = Id$  then there exists constant  $\beta > 0$  such that  $\mathbf{E}|u_j^{(k)}|^2 \le e^{2\beta jh} \mathbf{E}|u_0^{(k)}|^2 + 2K\gamma^2 \left(\frac{e^{2\beta jh} - 1}{e^{2\beta h} - 1}\right)$ 

**Rmk**. This becomes useless as  $h \rightarrow 0$ 

## For observations with $h \ll 1$ , we need another approach.

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## The EnKF equations look like a discretization

Recall the ensemble update equation

$$u_{j+1}^{(k)} = \Psi_h(u_j^{(k)}) + G(u_j) \left( \mathbf{y}_{j+1}^{(k)} - H \Psi_h(u_j^{(k)}) \right)$$
  
=  $\Psi_h(u_j^{(k)}) + \widehat{C}_{j+1} H^T (H^T \widehat{C}_{j+1} H + \Gamma)^{-1} \left( \mathbf{y}_{j+1}^{(k)} - H \Psi_h(u_j^{(k)}) \right)$ 

Subtract  $u_i^{(k)}$  from both sides and divide by h

$$\frac{u_{j+1}^{(k)} - u_{j}^{(k)}}{h} = \frac{\Psi_{h}(u_{j}^{(k)}) - u_{j}^{(k)}}{h} + \widehat{C}_{j+1}H^{T}(hH^{T}\widehat{C}_{j+1}H + h\Gamma)^{-1}\left(y_{j+1}^{(k)} - H\Psi_{h}(u_{j}^{(k)})\right)$$

Clearly we need to rescale the noise (ie.  $\Gamma$ ).

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#### Continuous-time limit

If we set  $\Gamma = h^{-1}\Gamma_0$  and substitute  $y_{j+1}^{(k)}$ , we obtain

$$\frac{u_{j+1}^{(k)} - u_j^{(k)}}{h} = \frac{\Psi_h(u_j^{(k)}) - u_j^{(k)}}{h} + \widehat{C}_{j+1}H^T(hH^T\widehat{C}_{j+1}H + \Gamma_0)^{-1} \\ \left(H_V + h^{-1/2}\Gamma_0^{1/2}\xi_{j+1} + h^{-1/2}\Gamma_0^{1/2}\xi_{j+1}^{(k)} - H\Psi_h(u_j^{(k)})\right)$$

But we know that

$$\Psi_h(\boldsymbol{u}_j^{(k)}) = \boldsymbol{u}_j^{(k)} + O(h)$$

and

$$\begin{split} \widehat{C}_{j+1} &= \frac{1}{K} \sum_{k=1}^{K} (\Psi_h(u_j^{(k)}) - \overline{\Psi_h(u_j)})^T (\Psi_h(u_j^{(k)}) - \overline{\Psi_h(u_j)}) \\ &= \frac{1}{K} \sum_{k=1}^{K} (u_j^{(k)} - \overline{u_j})^T (u_j^{(k)} - \overline{u_j}) + O(h) = C(u_j) + O(h) \end{split}$$

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## Continuous-time limit

We end up with

$$\frac{u_{j+1}^{(k)} - u_j^{(k)}}{h} = \frac{\Psi_h(u_j^{(k)}) - u_j^{(k)}}{h} - C(u_j)H^T\Gamma_0^{-1}H(u_j^{(k)} - v_j) + C(u_j)H^T\Gamma_0^{-1}\left(h^{-1/2}\xi_{j+1} + h^{-1/2}\xi_{j+1}^{(k)}\right) + O(h)$$

This looks like a numerical scheme for Itô S(P)DE

$$\frac{du^{(k)}}{dt} = F(u^{(k)}) - C(u)H^{T}\Gamma_{0}^{-1}H(u^{(k)} - v) \qquad (\bullet) \\ + C(u)H^{T}\Gamma_{0}^{-1/2}\left(\frac{dW^{(k)}}{dt} + \frac{dB}{dt}\right) .$$

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Nudging

$$\frac{du^{(k)}}{dt} = F(u^{(k)}) - C(u)H^{T}\Gamma_{0}^{-1}H(u^{(k)} - v) \qquad (\bullet) \\ + C(u)H^{T}\Gamma_{0}^{-1/2}\left(\frac{dW^{(k)}}{dt} + \frac{dB}{dt}\right) .$$

- 1 Extra dissipation term only sees differences in observed space
- ${\bf 2}$  Extra dissipation only occurs in the space spanned by ensemble

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## Kalman-Bucy limit

If F were **linear** and we write  $m(t) = \frac{1}{K} \sum_{k=1}^{K} \frac{u^{(k)}(t)}{k}$  then

$$\frac{dm}{dt} = F(m) - C(u)H^{T}\Gamma_{0}^{-1}H(m-v) + C(u)H^{T}\Gamma_{0}^{-1/2}\frac{dB}{dt} + O(K^{-1/2}).$$

This is the equation for the **Kalman-Bucy** filter, with empirical covariance C(u). The remainder  $O(K^{-1/2})$  can be thought of as a **sampling error**.

## Continuous-time results

Theorem (AS,DK)

Suppose that  $\{u^{(k)}\}_{k=1}^{K}$  satisfy  $(\bullet)$  with  $H = \Gamma = Id$ . Let

 $\mathbf{e}^{(k)} = \mathbf{u}^{(k)} - \mathbf{v} \; .$ 

Then there exists constant  $\beta > 0$  such that

$$\frac{1}{\mathcal{K}}\sum_{k=1}^{\mathcal{K}} \mathbf{E} |\boldsymbol{e}^{(k)}(t)|^2 \leq \left(\frac{1}{\mathcal{K}}\sum_{k=1}^{\mathcal{K}} \mathbf{E} |\boldsymbol{e}^{(k)}(0)|^2\right) \exp\left(\beta t\right) \ .$$

Why do we need  $H = \Gamma = Id$  ?

In the equation

$$\begin{aligned} \frac{d u^{(k)}}{dt} &= F(u^{(k)}) - C(u) H^T \Gamma_0^{-1} H(u^{(k)} - v) \\ &+ C(u) H^T \Gamma_0^{-1/2} \left( \frac{d W^{(k)}}{dt} + \frac{d B}{dt} \right) \;. \end{aligned}$$

The **energy** pumped in by the noise must be balanced by **contraction** of  $(u^{(k)} - v)$ . So the operator

$$C(\mathbf{u})H^{T}\Gamma_{0}^{-1}H$$

must be positive-definite.

Both C(u) and  $H^T \Gamma_0^{-1} H$  are pos-def, but this doesn't guarantee the same for the **product**!

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(1) Writing down an SDE/SPDE allows us to see the **important quantities** in the algorithm.

(2) Does not "prove" that catastrophic filter divergence is a numerical phenomenon, but is a decent starting point.

(1) Improve the condition on H.

(2) If we can **measure** the important quantities, then we can test the performance during the algorithm.