

# Data assimilation in high dimensions

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## What is data assimilation?

Suppose  $u$  satisfies

$$\frac{du}{dt} = F(u)$$

with some **unknown** initial condition  $u_0$ . We are most interested in geophysical models, so think high dimensional, nonlinear, possibly stochastic.

Suppose we make *partial, noisy* observations at times  $t = h, 2h, \dots, nh, \dots$

$$y_n = Hu_n + \xi_n$$

where  $H$  is a linear operator (think low rank projection),  $u_n = u(nh)$ , and  $\xi_n \sim N(0, \Gamma)$  iid.

The aim of **data assimilation** is to say something about the conditional distribution of  $u_n$  given the observations  $\{y_1, \dots, y_n\}$

## How does filtering work: (initialization)

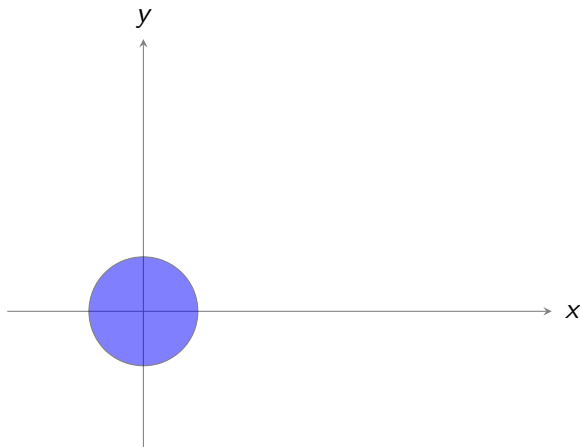
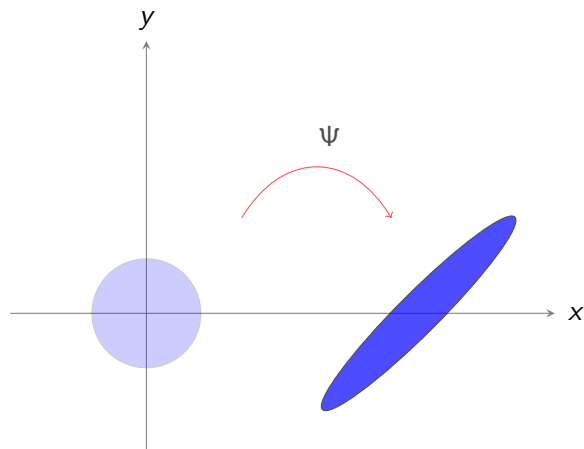


Figure: The blue circle represents our guess of  $u_0$ . Due to the uncertainty in  $u_0$ , this is a probability measure.

## How does filtering work: (forecast)



**Figure:** Apply the time  $h$  flow map  $\Psi$ . This produces a new probability measure which is our forecasted estimate of  $u_1$ . This is called the forecast step.

## How does filtering work: (make an observation)

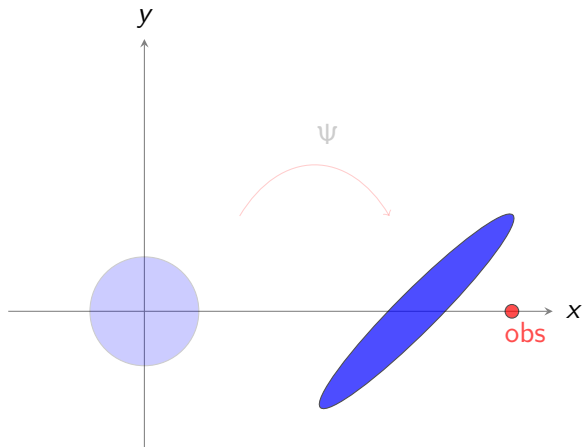


Figure: We make an observation  
 $y_1 = H u_1 + \xi_1$ . In the picture, we only observe the  $x$  variable.

## How does filtering work: (find best fit using Bayes)

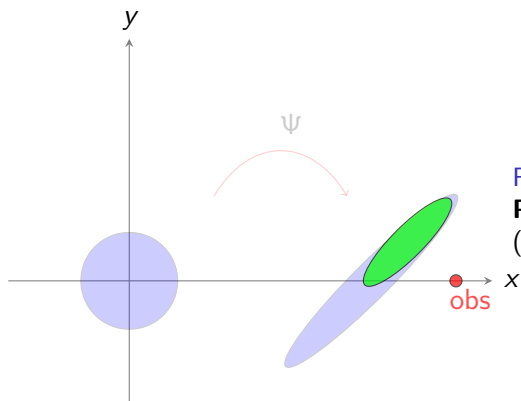


Figure:

$$\mathbf{P}(u_1|y_1) \propto \mathbf{P}(y_1|u_1)\mathbf{P}(u_1)$$

(Bayes formula)

## Problems in high dimensions

In **numerical weather prediction** the state dimension is  $O(10^9)$ .

- 1) Difficult to store a density of this size
- 2) Computing the 'forecast step' is an integration over the state space.

We need **low dimensional** approximations of the filtering problem.

We will look at the **Ensemble Kalman filter**.

## Ensemble Kalman filter (Evensen 94)

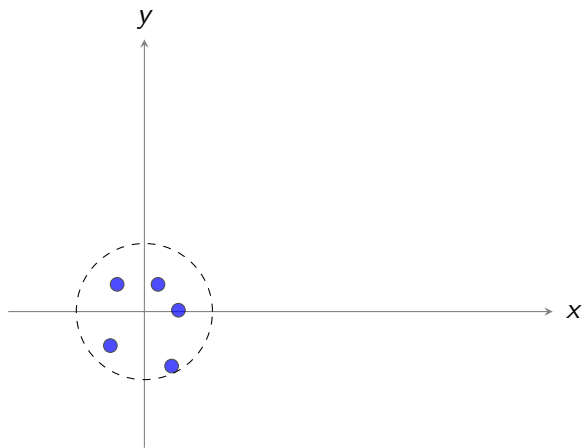


Figure: Start with  $K$  ensemble members drawn from some distribution. Empirical representation of  $u_0$ . The ensemble members are denoted  $v_0^{(k)}$ .

Only  $KN$  numbers are stored.



## Ensemble Kalman filter (Forecast step)

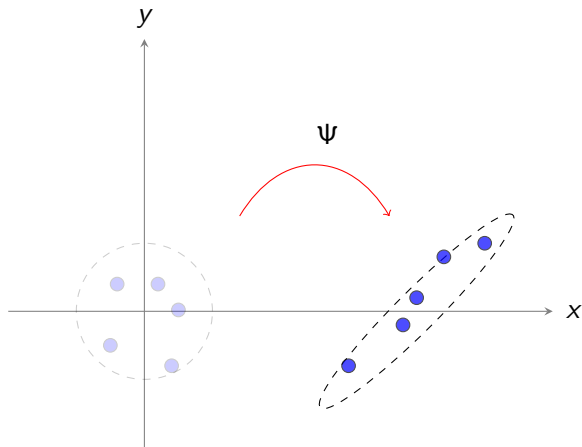


Figure: Apply the dynamics  $\Psi$  to each ensemble member.

$$v_0^{(k)} \mapsto \Psi(v_0^{(k)})$$

## Ensemble Kalman filter (Make obs)

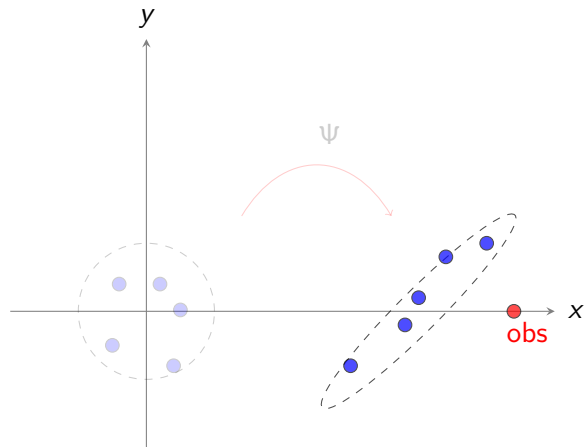


Figure: Make an observation.

## Ensemble Kalman filter (Perturb obs)

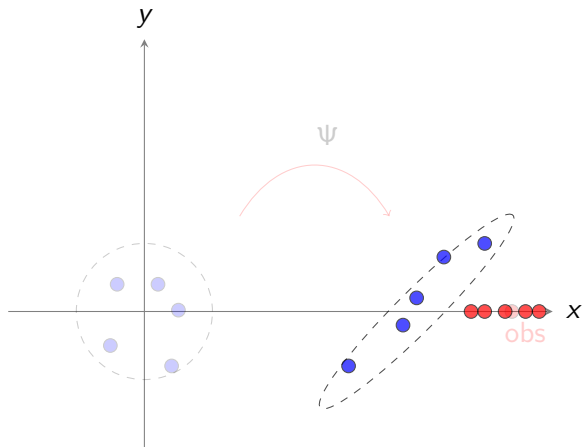


Figure: Turn the observation into  $K$  artificial observations by perturbing by the same source of observational noise.

$$y_1^{(k)} = y_1 + \xi_1^{(k)}$$

## Ensemble Kalman filter (find best fit using Bayes)

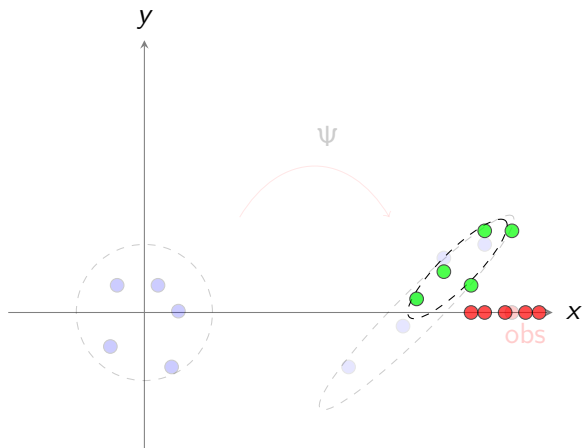


Figure: Update each member using the 'Kalman update formula'. This is a linear approximation of Bayes.

$$v_1^{(k)} = \Psi(v_0^{(k)}) + K_1(y_1^{(k)} - H\Psi(v_0^{(k)}))$$

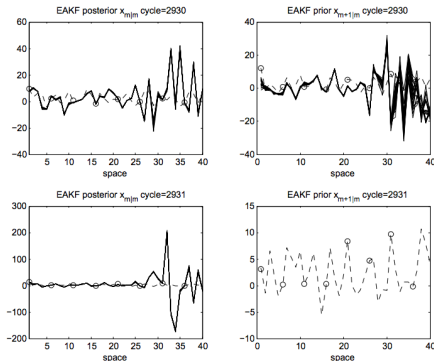
## Why should mathematicians be interested?

A widely used algorithm (NWP, disease forecasting, chemical reactions) with many questions and not so many answers:

- 1 - Is the filter stable to perturbations? *eg. Will different initializations converge? (ergodicity)*
- 2 - Is the filter accurate? *Is the posterior consistent with the true signal?*
- 3 - Can we design mathematically sensible alternative algorithms?
- 4 - Can we understand why/when the filter fails?

## Catastrophic filter divergence

Lorenz-96:  $\dot{u}_j = (u_{j+1} - u_{j-2})u_{j-1} - u_j + F$  with  $j = 1, \dots, 40$ . Periodic BCs. Observe every fifth node. (*Harlim-Majda 10, Gottwald-Majda 12*)



True solution in a bounded set, but filter **blows up** to machine infinity in finite time!

For complicated models, only heuristic arguments offered as explanation.

*Can we **prove** it for a simpler constructive model?*

## The rotate-and-lock map (K., Majda, Tong. PNAS 15.)

The model  $\Psi : \mathbb{R}^2 \rightarrow \mathbb{R}^2$  is a composition of two maps  $\Psi(x, y) = \Psi_{lock}(\Psi_{rot}(x, y))$  where

$$\Psi_{rot}(x, y) = \begin{pmatrix} \rho \cos \theta & -\rho \sin \theta \\ \rho \sin \theta & \rho \cos \theta \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix}$$

and  $\Psi_{lock}$  rounds the input to the nearest point in the grid

$$\mathcal{G} = \{(m, (2n + 1)\varepsilon) \in \mathbb{R}^2 : m, n \in \mathbb{Z}\}.$$

It is easy to show that this model has an **energy dissipation principle**:

$$|\Psi(x, y)|^2 \leq \alpha |(x, y)|^2 + \beta$$

for  $\alpha \in (0, 1)$  and  $\beta > 0$ .



(a)

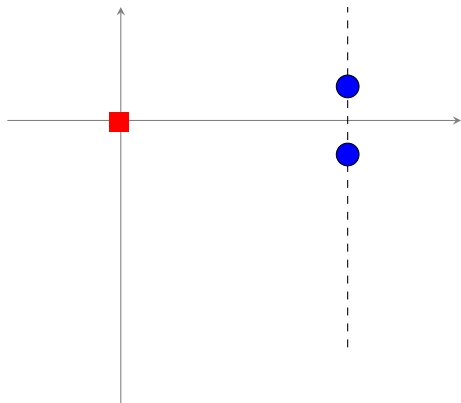


Figure: The red square is the trajectory  $u_n = 0$ . The blue dots are the positions of the forecast ensemble  $\Psi(v_0^+)$ ,  $\Psi(v_0^-)$ . Given the locking mechanism in  $\Psi$ , this is a natural configuration.

(b)

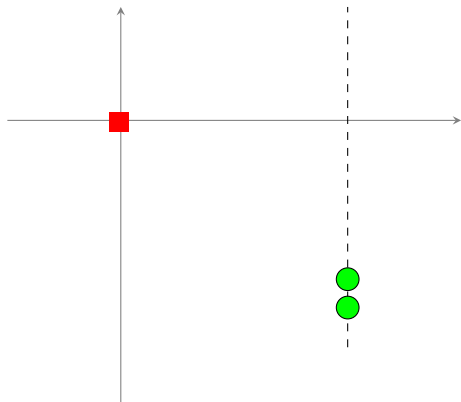


Figure: We make an observation ( $H$  shown below) and perform the analysis step. The green dots are  $v_1^+$ ,  $v_1^-$ .

Observation matrix  
$$H = \begin{pmatrix} 1 & 0 \\ \varepsilon^{-2} & 1 \end{pmatrix}$$

The filter is 'sure' that  $u_1 = \hat{x}$  (the dashed line). The filter deduces that the observation is approximately  $(y_1, y_2) = (\hat{x}, \varepsilon^{-2}\hat{x} + u_2)$ .  
Thus  $v_1^\pm \approx (\hat{x}, -\varepsilon^{-2}\hat{x})$

(c)

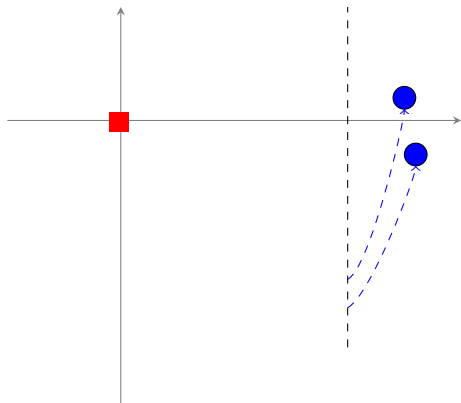


Figure: Beginning the next assimilation step. Apply  $\Psi_{rot}$  to the ensemble (blue dots)

(d)

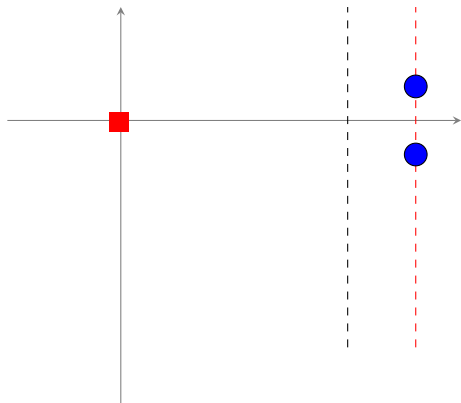


Figure: Apply  $\Psi_{lock}$ .  
The blue dots are the forecast ensemble  $\Psi(v_1^+)$ ,  $\Psi(v_1^-)$ . Exact same as frame 1, but higher energy orbit. The cycle repeats leading to **exponential growth**.

## Theorem (K.-Majda-Tong 15 PNAS)

*For any  $N > 0$  and any  $p \in (0, 1)$  there exists a choice of parameters such that*

$$\mathbf{P} \left( |\mathbf{v}_n^{(k)}| \geq M_n \text{ for all } n \leq N \right) \geq 1 - p$$

*where  $M_n$  is an exponentially growing sequence.*

**ie** - The filter can be made to grow exponentially for an arbitrarily long time with an arbitrarily high probability.

## References

- 1 - D. Kelly, K. Law & A. Stuart. *Well-Posedness And Accuracy Of The Ensemble Kalman Filter In Discrete And Continuous Time*. **Nonlinearity** (2014).
- 2 - D. Kelly, A. Majda & X. Tong. *Concrete ensemble Kalman filters with rigorous catastrophic filter divergence*. **Proc. Nat. Acad. Sci.** (2015).
- 3 - X. Tong, A. Majda & D. Kelly. *Nonlinear stability and ergodicity of ensemble based Kalman filters*. **Nonlinearity** (2016).
- 4 - X. Tong, A. Majda & D. Kelly. *Nonlinear stability of the ensemble Kalman filter with adaptive covariance inflation*. To appear in **Comm. Math. Sci.** (2015).

All my slides are on my website ([www.dtbkelly.com](http://www.dtbkelly.com)) **Thank you!**